# **Celestial Amplitudes and Light Transforms**

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The quest for flat space holography has recently received a boost owing to the realization that scattering amplitudes in 4D flat spacetime can be recast as correlation functions of a 2D conformal field theory living on the celestial sphere [1, 2]. Then the celestial CFT (CCFT) becomes a potential candidate for a holographic description of the flat space S-matrix. In generic CFTs, the OPE coefficients are related to three-point functions and four-point functions contain information about the spectrum of the theory, which can be deduced by means of the conformal block decomposition. In CCFT, three- and (treelevel) four-point correlators make these relationships opaque due to the distributional nature of their correlators. It was shown in [3] that certain three-point correlators involving light-ray operators take the form of standard three-point CFT correlators. In this poster, we

#### **Correlator of 2 light-ray gluons and 2 gluons** 3.

The correlator of two light-ray gluon operators and two gluon operators can be described by the Gauss hypergeometric functions. For example, consider  $z, \overline{z} \in [0, 1]$ ,  $\left\langle \bar{\mathbf{L}}[\mathcal{O}_{\Delta_1,-}](z_1,\bar{z}_1)\bar{\mathbf{L}}[\mathcal{O}_{\Delta_2,-}](z_2,\bar{z}_2)\mathcal{O}_{\Delta_3,+}(z_3,\bar{z}_3)\mathcal{O}_{\Delta_4,+}(z_4,\bar{z}_4)\right\rangle$  $= -\pi \,\delta(\beta) \, X\left(z_{ij}, \bar{z}_{ij}\right) \, \left(\frac{z}{1-z}\right) \, |z\bar{z}|^{1-\frac{\Delta_1+\Delta_2}{2}}$ (7)  $\left\{ |z-1|^{1-\Delta_3} {}_2F_1\left[ 1-\Delta_2, \Delta_3-1, \Delta_3+\Delta_4-2, \frac{z-\bar{z}}{z-1} \right] C(\Delta_3-1, \Delta_4-1) \right\}$  $+ |z-1|^{\Delta_4-2} |\bar{z}-z|^{\Delta_1+\Delta_2-1} {}_2F_1 \left[ 2-\Delta_4, \Delta_1, \Delta_1+\Delta_2, \frac{z-\bar{z}}{z-1} \right] C(\Delta_1-1, \Delta_2-1) \right\}$ 



- present the four-point correlator of two gluon light-ray operators and two gluon primaries from the four-gluon celestial amplitude in (2, 2) signature spacetime. The correlator is non-distributional and allows us to verify that light-ray operators appear in the OPE of two gluon primaries. We also carry out a conformal block decomposition of the terms involving the exchange of gluon operators.
- present the correlator of four gluon light-ray operators in celestial CFT. We find that it is described by Fox H-functions and generalized I-functions of multiple variables.

## Preliminary

**Celestial amplitudes in (2,2) signature** Any (non-zero) null four-vector  $p^{\mu}$  in (2,2) signature can be uniquely parameterized as

$$\rho^{\mu} = \epsilon \omega \left( 1 + z\bar{z}, z + \bar{z}, z - \bar{z}, 1 - z\bar{z} \right) , \qquad (1)$$

where  $\epsilon = \pm 1$ ,  $\omega > 0$ , and z and  $\overline{z}$  are independent real variables.

- In (1,3) signature:  $\epsilon$  would indicate whether  $p^{\mu}$  describes an incoming or outgoing particle.
- In (2,2) signature:  $\epsilon$  labels different Poincaré patches.

Note that it transforms covariantly under  $SL(2,\mathbb{R}) \times SL(2,\mathbb{R})$  conformal transformations:

where C(a, b) = B(a, b) + B(a, 1 - a - b) + B(b, 1 - a - b),  $B(a, b) = \frac{\Gamma(a)\Gamma(b)}{\Gamma(a+b)}$ . Note that the  $|\bar{z} - z|^{\Delta_1 + \Delta_2 - 1}$  term contributes a bulk point singularity.

**Leading OPE structures via collinear limit** Note that in (7) there is no  $\delta(z - \bar{z})$  and we can read off the leading OPE structures via evaluating the collinear limits. For example, we can read off the two-gluonoperator-OPE as below, which matches with [4].

$$\mathcal{O}_{\Delta_{3},+}(z_{3},\bar{z}_{3})\mathcal{O}_{\Delta_{4},+}(z_{4},\bar{z}_{4}) \sim \frac{C(\Delta_{3}-1,\Delta_{4}-1)}{z_{34}}\mathcal{O}_{\Delta_{3}+\Delta_{4}-1,+}(z_{4},\bar{z}_{4}) + \frac{\bar{L}[\mathcal{O}_{\Delta_{3}+\Delta_{4}-1,+}](z_{4},\bar{z}_{4})}{z_{34}|\bar{z}_{34}|^{\Delta_{3}+\Delta_{4}-3}}$$
(8)

**Conformal Block Decompositions** For the violet term in (7), we compute its conformal block decomposition as

violet term = 
$$\sum_{n=0}^{\infty} \sum_{k=n}^{\infty} a_{k,n} k_{34}^{21} \left[ \frac{\Delta_3 + \Delta_4}{2} + k, \frac{\Delta_3 + \Delta_4}{2} + n - 1 \right]$$
, (9)

where the coefficients are

$$a_{k,n} = C(\Delta_3 - 1, \Delta_4 - 1) \frac{(1 - \Delta_1)_n (1 - \Delta_2)_n (\Delta_3 - 1)_n (\Delta_4 - 1)_n}{n! (\Delta_3 + \Delta_4 + n - 3)_n (\Delta_3 + \Delta_4 - 2)_{2n}} \times \sum_{m=0}^{k-n} \frac{(1 - \Delta_2 + n)_m (\Delta_4 - 1 + n)_m}{(\Delta_3 + \Delta_4 + 2n - 2)_{2m}} \frac{(2 - \Delta_1 + n + m)_{k-n-m} (\Delta_3 + n + m)_{k-n-m}}{(\Delta_3 + \Delta_4 - 1 + 2n + 2m)_{2k-2n-2m}}$$
(10)

and  $k_{34}^{21}[h, \bar{h}]$  are the usual conformal blocks [5]

$$x_{34}^{21}[h,\bar{h}] = {}_{2}F_{1}[h-h_{12},h+h_{34},2h,z]\bar{z}^{\bar{h}-\bar{h}_{3}-\bar{h}_{4}}{}_{2}F_{1}[\bar{h}-\bar{h}_{12},\bar{h}+\bar{h}_{34},2\bar{h},\bar{z}].$$
(11)

$$\epsilon \to \epsilon \operatorname{sgn}((cz+d)(\bar{c}\bar{z}+\bar{d})) \ , \ z \to (az+b)/(cz+d) \ , \ \bar{z} \to (\bar{a}\bar{z}+\bar{b})/(\bar{c}\bar{z}+\bar{d})$$
(2)

Celestial and momentum space amplitudes for massless particles are related to each other by a change of basis provided by the Mellin transform:

$$\left\langle \mathcal{O}_{\Delta_{1},J_{1}}^{\epsilon_{1}}(z_{1},\bar{z}_{1})\cdots\mathcal{O}_{\Delta_{n},J_{n}}^{\epsilon_{n}}(z_{n},\bar{z}_{n})\right\rangle = \left(\prod_{i=1}^{n}\int_{0}^{\infty}d\omega_{i}\,\omega_{i}^{\Delta_{i}-1}\right)\,\mathcal{A}_{n}(\epsilon_{i},\omega_{i},z_{i},\bar{z}_{i}) \tag{3}$$

**Correlator of 4 gluon operators** The tree-level, color-ordered, four-gluon amplitude is given by the Parke-Taylor formula and its corresponding celestial amplitude obtained by Mellin transform is

$$\left\langle \mathcal{O}_{\Delta_{1},-}^{\epsilon_{1}}(z_{1},\bar{z}_{1})\mathcal{O}_{\Delta_{2},-}^{\epsilon_{2}}(z_{2},\bar{z}_{2})\mathcal{O}_{\Delta_{3},+}^{\epsilon_{3}}(z_{3},\bar{z}_{3})\mathcal{O}_{\Delta_{4},+}^{\epsilon_{4}}(z_{4},\bar{z}_{4})\right\rangle = \pi\,\delta(\beta)\,\delta(z-\bar{z})\,X\left(z_{ij},\bar{z}_{ij}\right) \\ \times \,\operatorname{sgn}\left(\frac{z}{z-1}\right)|z|^{3}|1-z|^{1-\Delta_{2}-\Delta_{3}}\,\Theta\left(-\epsilon_{1}\epsilon_{4}\frac{z_{24}\bar{z}_{24}}{z_{12}\bar{z}_{12}}z\right)\,\Theta\left(\epsilon_{2}\epsilon_{4}\frac{z_{34}\bar{z}_{34}}{z_{23}\bar{z}_{23}}\frac{1-z}{z}\right)\,\Theta\left(\epsilon_{3}\epsilon_{4}\frac{z_{24}\bar{z}_{24}}{z_{23}\bar{z}_{23}}(z-1)\right)$$

$$\text{where } X(z_{ij},\bar{z}_{ij}) = \frac{1}{|z_{12}|^{h_{1}+h_{2}}|z_{34}|^{h_{3}+h_{4}}}\left|\frac{z_{24}}{z_{14}}\right|^{h_{12}}\left|\frac{z_{14}}{z_{14}}\right|^{h_{34}}\frac{1}{|\bar{z}_{12}|^{\bar{h}_{1}+\bar{h}_{2}}|\bar{z}_{34}|^{\bar{h}_{3}+\bar{h}_{4}}}\left|\frac{\bar{z}_{24}}{\bar{z}_{14}}\right|^{\bar{h}_{34}}, h_{i} = (\Delta_{i}+J_{i})/2, \\ \bar{h}_{i} = (\Delta_{i}-J_{i})/2, h_{ij} = h_{i}-h_{j}, \bar{h}_{ij} = \bar{h}_{i}-\bar{h}_{j}, \text{ the cross ratio } z = (z_{12}z_{34})/(z_{13}z_{24}), \text{ and } \Theta(x) \text{ denotes the Heaviside step function.}$$

**Light transforms** The definitions of the light transforms are summarized as below.

Light Transform	Definition	<b>Conformal Weight</b>
Anti-holomorphic	$\bar{\mathbf{L}}[\mathcal{O}_{h,\bar{h}}](z,\bar{z}) = \int_{\mathbb{R}} \frac{d\bar{z}'}{ \bar{z}'-\bar{z} ^{2-2\bar{h}}} \mathcal{O}_{h,\bar{h}}(z,\bar{z}')$	$(h, 1 - \overline{h})$
Holomorphic	$\mathbf{L}[\mathcal{O}_{h,\bar{h}}](z,\bar{z}) = \int_{\mathbb{R}} \frac{dz'}{ z'-z ^{2-2h}} \mathcal{O}_{h,\bar{h}}(z',\bar{z})$	$(1-h, \bar{h})$

**Step functions & spacetime signatures** The step functions appearing in (4) make light transform in-

#### . Correlator of 4 light-ray gluon operators 4.

We consider the correlator involving two holomorphic and two anti-holomorphic light-ray operators:

$$\left\langle \bar{\mathbf{L}}[\mathcal{O}_{\Delta_{1},-}](z_{1},\bar{z}_{1})\,\bar{\mathbf{L}}[\mathcal{O}_{\Delta_{2},-}](z_{2},\bar{z}_{2})\,\mathbf{L}[\mathcal{O}_{\Delta_{3},+}](z_{3},\bar{z}_{3})\,\mathbf{L}[\mathcal{O}_{\Delta_{4},+}](z_{4},\bar{z}_{4})\right\rangle$$

$$= \pi\,\delta(\beta)\,X\left(z_{ij},\bar{z}_{ij}\right)\,\left|\frac{z}{\bar{z}}\right|^{\frac{\Delta_{1}+\Delta_{2}-2}{2}}|1-z|^{2-\Delta_{2}-\Delta_{4}}\,\mathcal{G}(z,\bar{z})$$
(12)

where  $\mathcal{G}(z, \overline{z})$  function can be organized as

$$\mathcal{G}(z,\bar{z}) = \int_{-\infty}^{\infty} \frac{dx}{x(x-1)} \mathcal{F}_1(z,x) \mathcal{F}_2(x,\bar{z}) \sim \int_{-\infty}^{\infty} dx \left(\prod_i |x-a_i|^{\alpha_i}\right) {}_2F_1(\cdots) {}_2F_1(\cdots)$$
(13)

Here, both  $\mathcal{F}_1(z, x)$  and  $\mathcal{F}_2(x, \overline{z})$  are four-marked-point integrals

$$\mathcal{F}_1(z,x) = \int_{-\infty}^{\infty} dy \, |y-z|^{\Delta_4 - 1} |x-y|^{\Delta_3 - 1} |y-1|^{\Delta_2 - 2} \,, \tag{14}$$

$$\mathcal{F}_2(x,\bar{z}) = \int_{-\infty}^{\infty} dt \, |x-t|^{\Delta_1 - 1} |1-t|^{\Delta_4 - 2} |\bar{z}-t|^{\Delta_2 - 1} \,. \tag{15}$$

and basically Gauss Hypergeometric functions. Our strategy to evaluate the  $\mathcal{G}(z, \overline{z})$  function: 1. Use Mellin-Barnes representation of the Gauss hypergeometric function to extract the x-integral

$$= \Gamma - \frac{1}{1} - \frac{\Gamma(c)}{1} - \frac{\Gamma(a)}{1} - \frac{\Gamma(-s)\Gamma(a+s)\Gamma(b+s)}{1} - \frac{\Gamma(-s)\Gamma(a+s)\Gamma(b+s)}{1} - \frac{\Gamma(a+s)\Gamma(b+s)}{1} - \frac{\Gamma(a+s)\Gamma(b+s)}{1}$$

$${}_{2}F_{1}[a,b,c;-|z|] = \frac{1}{2\pi i} \frac{\Gamma(c)}{\Gamma(a)\Gamma(b)} \int_{-i\infty}^{i\infty} ds \frac{\Gamma(-s)\Gamma(a+s)\Gamma(b+s)}{\Gamma(c+s)} |z|^{s}$$
(16)

2. Do the *x*-integral first

3. Realize the remaining multivariable Mellin-Barnes-type integrals as the Fox H-function H(x, y) or Generalized I-functions  $I(z_1, \ldots, z_r)$ .

$$\mathbf{H}(x,y) = \frac{1}{(2\pi i)^2} \int_{-i\infty}^{+i\infty} ds \int_{-i\infty}^{+i\infty} dt \,\phi_1(s+t) \,\phi_2(s) \,\phi_3(t) \, x^{-s} \, y^{-t} \tag{17}$$

tegrals complicated due to the independence of z and  $\overline{z}$  in (2,2) spacetime.

$(1,3) \rightarrow \text{Euclidean CFT}$	$(2,2) \rightarrow \text{Lorentzian CFT}$
is the complex conjugate of $\boldsymbol{z}$	$z$ and $ar{z} \in \mathbb{R}$ are independent
	$sgn(z_{ij}\bar{z}_{ij}) > 0$ : spacelike
$\operatorname{sgn}(z_{ij}\bar{z}_{ij}) \ge 0$	$sgn(z_{ij}\bar{z}_{ij}) = 0$ : null-separated
	$sgn(z_{ij}\bar{z}_{ij}) < 0:timelike$

Our resolution is based on the fact that

$$\sum_{=\pm} \Theta \left( -\epsilon_1 \epsilon_4 \frac{z_{24} \bar{z}_{24}}{z_{12} \bar{z}_{12}} z \right) \Theta \left( \epsilon_2 \epsilon_4 \frac{z_{34} \bar{z}_{34}}{z_{23} \bar{z}_{23}} \frac{1-z}{z} \right) \Theta \left( \epsilon_3 \epsilon_4 \frac{z_{24} \bar{z}_{24}}{z_{23} \bar{z}_{23}} (z-1) \right) = 2 ,$$
(5)

we define the following object and we will study the light transforms of it.

 $\left\langle \mathcal{O}_{\Delta_{1},-}(z_{1},\bar{z}_{1})\mathcal{O}_{\Delta_{2},-}(z_{2},\bar{z}_{2})\mathcal{O}_{\Delta_{3},+}(z_{3},\bar{z}_{3})\mathcal{O}_{\Delta_{4},+}(z_{4},\bar{z}_{4})\right\rangle := \sum_{\epsilon:=+}\left\langle \mathcal{O}_{\Delta_{1},-}^{\epsilon_{1}}(z_{1},\bar{z}_{1})\mathcal{O}_{\Delta_{2},-}^{\epsilon_{2}}(z_{2},\bar{z}_{2})\mathcal{O}_{\Delta_{3},+}^{\epsilon_{3}}(z_{3},\bar{z}_{3})\mathcal{O}_{\Delta_{4},+}^{\epsilon_{4}}(z_{4},\bar{z}_{4})\right\rangle \quad \text{(6)}$ 

 $\mathbf{I}(z_1,\ldots,z_r) = \frac{1}{(2\pi i)^r} \int_{-i\infty}^{+i\infty} \ldots \int_{-i\infty}^{+i\infty} ds_1 \ldots ds_r \,\phi(s_1,\ldots,s_r) \,\theta_1(s_1) \ldots \theta_r(s_r) \, z_1^{s_1} \ldots z_r^{s_r}$ (18)

### References

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